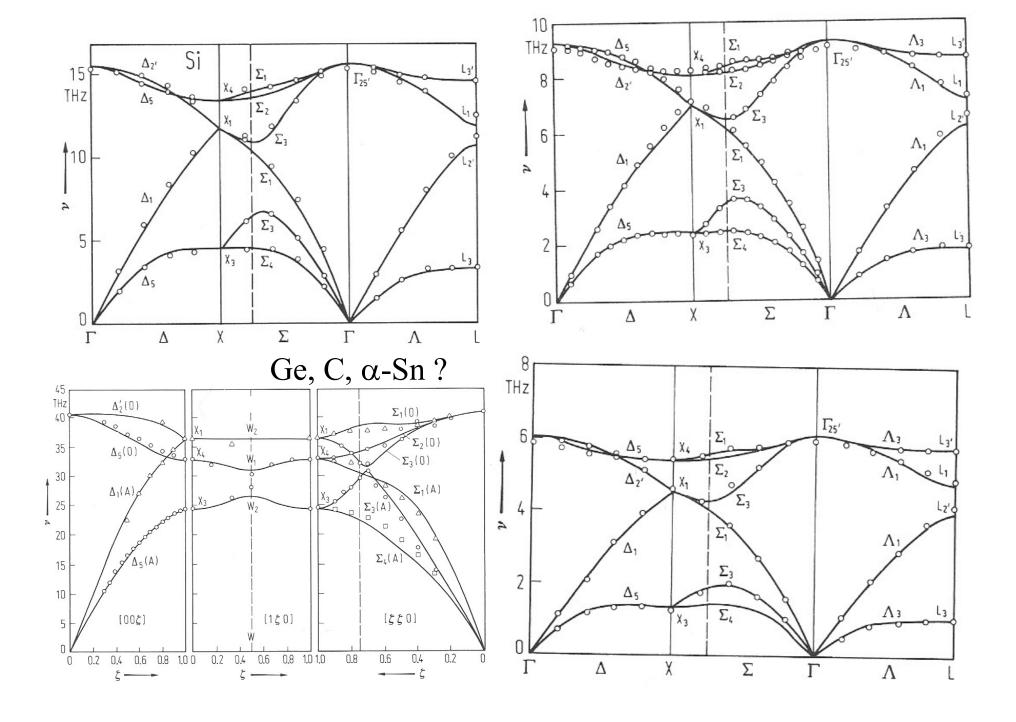


Technische Universität Graz

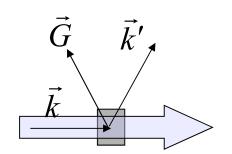
## Phonons



#### Inelastic neutron scattering

Diffraction condition for elastic scattering

$$\vec{k}' = \vec{k} + \vec{G}$$



The whole crystal recoils with momentum

$$\hbar \vec{G}$$

Diffraction condition for inelastic scattering

$$\vec{k}' \pm \vec{K}_{ph} = \vec{k} + \vec{G}$$
  $\frac{\hbar^2 k'^2}{2m_n} \pm \hbar \omega_{ph} = \frac{\hbar^2 k^2}{2m_n} + \frac{\hbar^2 G^2}{2m_{crystal}}$ 

 $\vec{K}_{ph}$  is the phonon momentum

Phonon dispersion relations are determined experimentally by inelastic neutron diffraction

#### long wavelength limit

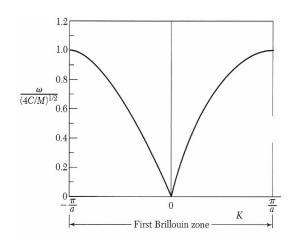
discrete version of wave equation

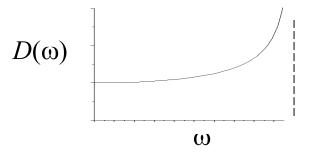
$$m\frac{d^2u_s}{dt^2} = C(u_{s+1} - 2u_s + u_{s-1})$$

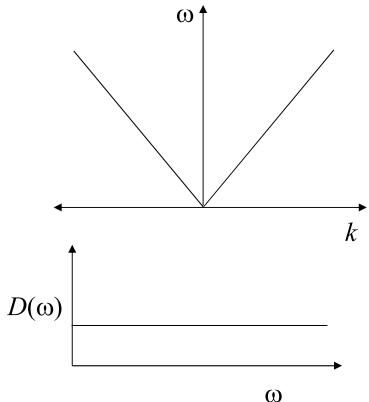
1-d wave equation

$$\frac{d^2u}{dt^2} = c^2 \frac{d^2u}{dx^2}$$

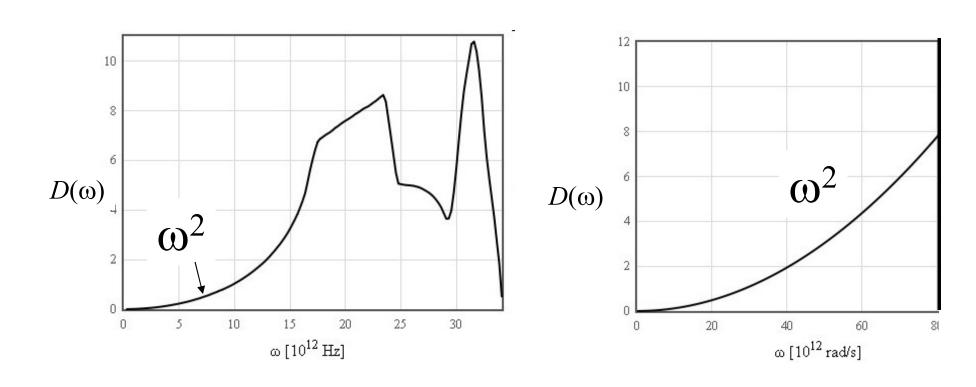
The solutions to the linear chain are the same as the solutions to the wave equation for  $|k| << \pi/a$ .







# long wavelength, low temperature limit



#### Phonons - long wavelength, low temperature limit

$$I = \frac{2\pi^5 k_B^4 T^4}{15c^2 h^3} = \sigma T^4 \qquad [\text{J m}^{-2} \text{ s}^{-2}]$$

At low T, there are only long wave length states occupied.

3 polarizations

Density of states per polarization:

$$D(\omega)d\omega = \frac{\omega^2}{c^3\pi^2}d\omega.$$

Specific heat of insulators at low temperatures per polarization

$$c_{\nu} = \frac{816\sigma T^3}{c}$$
 [J K<sup>-1</sup> m<sup>-3</sup>]  $f = \frac{-4\sigma T^4}{3c}$  [J/m<sup>3</sup>]



Speed of sound

$$u(\lambda) = \frac{8\pi hc}{\lambda^{5} \left( \exp\left(\frac{hc}{\lambda k_{B}T}\right) - 1 \right)} \qquad [J/m^{4}]$$

$$u = \frac{4\sigma T^4}{c} \qquad [J/m^3]$$

$$c_{\nu} = \frac{16\sigma T^3}{c} \qquad [\text{J K}^{-1} \text{ m}^{-3}]$$

$$f = \frac{-4\sigma T^4}{3c} \qquad [J/m^3]$$

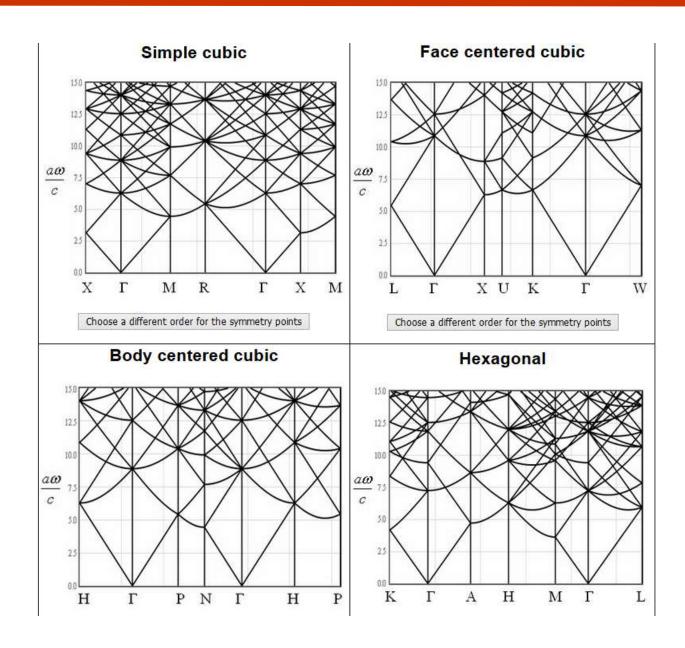
$$s = \frac{16\sigma T^3}{3c}$$
 [J K<sup>-1</sup> m<sup>-3</sup>]

$$P = \frac{4\sigma T^4}{3c} \qquad [N/m^2]$$

#### Empty lattice approximation

Use the speed of sound instead of the speed of light.

3 acoustic branches 3*p* - 3 optical branches



#### Thermal properties

#### 1. Determine the dispersion relation:

Write down the equations of motion (masses and springs).

The solutions to these equations will be eigen functions of T

$$\exp\left(i\left(\vec{k}\cdot\vec{a}_1 + \vec{k}\cdot\vec{a}_2 + \vec{k}\cdot\vec{a}_3 - \omega t\right)\right)$$

Substitute the eigen functions of T into the equations of motion to determine the dispersion relation.

## 2. Determine the density of states numerically from the dispersion relation $D(\omega)$

For every allowed k, find all corresponding values of  $\omega$ .

#### Specific Heat

$$u(T) = \int\limits_0^\infty rac{\hbar \omega D(\omega)}{\exp\left(rac{\hbar \omega}{k_B T}
ight) - 1} \, d\omega$$

$$c_v = \left(rac{\partial u}{\partial T}
ight)_{N,V}$$

$$c_v = \int \hbar \omega D(\omega) rac{\partial}{\partial T} \left(rac{1}{e^{rac{\hbar \omega}{k_B T}} - 1}
ight) d\omega$$

$$c_v = \int \left(rac{\hbar\omega}{T}
ight)^2 rac{D(\omega)e^{rac{\hbar\omega}{k_BT}}}{k_Bigg(e^{rac{\hbar\omega}{k_BT}}-1igg)^2}d\omega$$

http://lampx.tugraz.at/~hadley/ss1/phonons/table/dos2cv.html

#### Heat capacity / specific heat

**Heat capacity** is the measure of the heat energy required to increase the temperature of an object by a certain temperature interval.

**Specific heat** is the measure of the heat energy required to increase the temperature of a unit quantity of a substance by a certain temperature interval.

For solids, the heat capacity at constant volume and heat capacity at constant pressure are almost the same.

The heat capacity was historically important for understanding solids.

## Dulong and Petit (Classical result)

Equipartition:  $\frac{1}{2}k_BT$  per quadratic term in energy

$$u = 3nk_{\scriptscriptstyle R}T$$

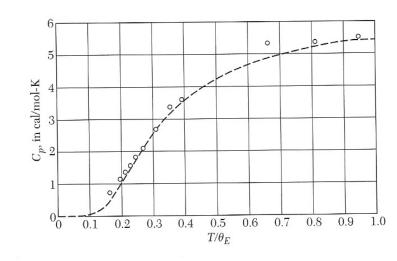
internal energy:  $u = 3nk_BT$  n = atomic density

specific heat: 
$$c_v = \frac{du}{dT} = 3nk_B$$

experiments: heat capacity goes to zero at zero temperature



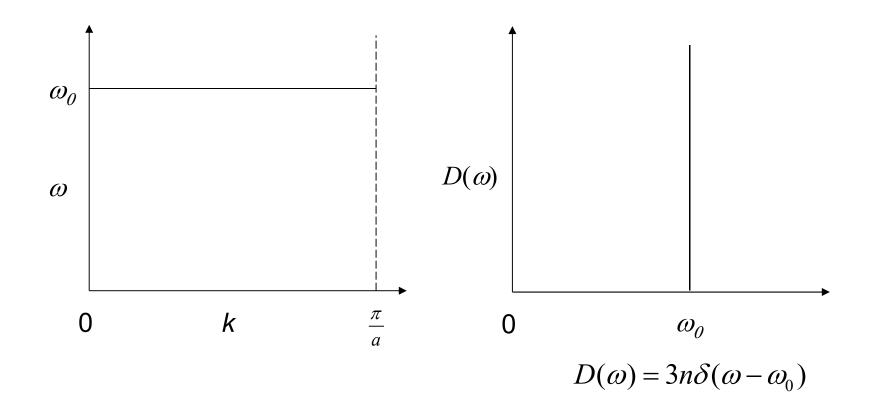
Pierre Louis Dulong





Alexis Therese Petit

#### Einstein model for specific heat



n =density of atoms

$$u(\omega) = D(\omega)\hbar\omega \frac{1}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1} = \hbar\omega \frac{3n\delta(\omega - \omega_0)}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1}$$

#### Einstein model for specific heat

$$u(\omega) = \hbar \omega \frac{3n\delta(\omega - \omega_0)}{\exp\left(\frac{\hbar \omega}{k_B T}\right) - 1}$$

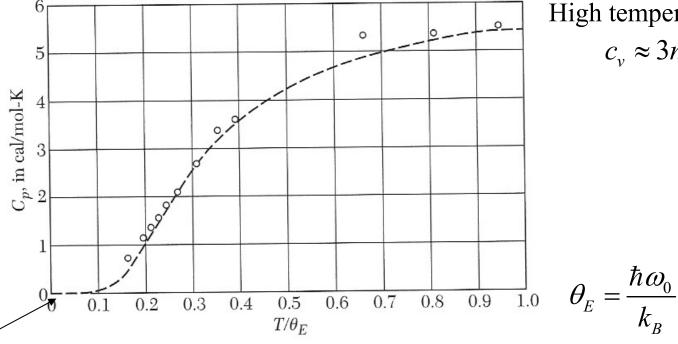
$$\frac{u(T)}{3n\hbar\omega_{0}} = \frac{u(T)}{0.1} =$$

$$u = \int_{0}^{\infty} u(\omega)d\omega = \int_{0}^{\infty} 3n\hbar\omega \frac{\delta(\omega - \omega_{0})d\omega}{\exp\left(\frac{\hbar\omega}{k_{B}T}\right) - 1} = \frac{3n\hbar\omega_{0}}{\exp\left(\frac{\hbar\omega_{0}}{k_{B}T}\right) - 1}$$

$$c_{v} = \frac{du}{dT} = \frac{3n\hbar\omega_{0} \frac{\hbar\omega_{0}}{k_{B}T^{2}} \exp\left(\frac{\hbar\omega_{0}}{k_{B}T}\right)}{\left(\exp\left(\frac{\hbar\omega_{0}}{k_{B}T}\right)^{-1}\right)^{2}} = \frac{3nk_{B}\left(\frac{\hbar\omega_{0}}{k_{B}T}\right)^{2} \exp\left(\frac{\hbar\omega_{0}}{k_{B}T}\right)}{\left(\exp\left(\frac{\hbar\omega_{0}}{k_{B}T}\right) - 1\right)^{2}}$$

### Einstein model for specific heat

$$c_{v} = \frac{3nk_{B} \left(\frac{\hbar\omega_{0}}{k_{B}T}\right)^{2} \exp\left(\frac{\hbar\omega_{0}}{k_{B}T}\right)}{\left(\exp\left(\frac{\hbar\omega_{0}}{k_{B}T}\right) - 1\right)^{2}}$$



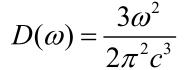
High temperatures  $c_v \approx 3nk_B$ 

$$\theta_E = \frac{\hbar \omega_0}{k_B}$$

low T does not fit

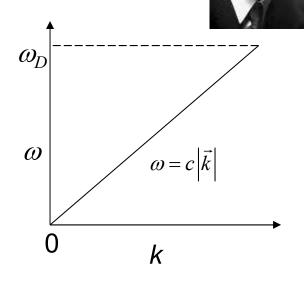
## Debye model for specific heat





Like blackbody radiation up to a cutoff frequency.

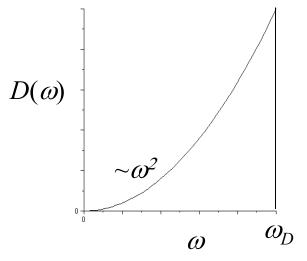
$$V \int_{0}^{\infty} D(\omega) d\omega = 3N = Na^{3} \int_{0}^{\omega_{D}} \frac{3\omega^{2}}{2\pi^{2}c^{3}} d\omega = Na^{3} \frac{\omega_{D}^{3}}{2\pi^{2}c^{3}}$$



$$\omega_D = \left(\frac{6\pi^2 c^3}{a^3}\right)^{1/3}$$

Debye temperature

$$\hbar\omega_D = k_B \theta_D$$



#### Debye model for heat capacity

$$u(\omega) = D(\omega)\hbar\omega \frac{1}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1} = \frac{3\omega^2}{2\pi^2 c^3} \frac{\hbar\omega}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1}$$

$$u = \int_{0}^{\omega_{D}} u(\omega)d\omega = \int_{0}^{\omega_{D}} \frac{3\omega^{2}}{2\pi^{2}c^{3}} \frac{\hbar\omega}{\exp\left(\frac{\hbar\omega}{k_{B}T}\right) - 1} d\omega \approx \int_{0}^{\infty} \frac{3\omega^{2}}{2\pi^{2}c^{3}} \frac{\hbar\omega}{\exp\left(\frac{\hbar\omega}{k_{B}T}\right) - 1} d\omega$$
for low  $T$ 

$$\int_{0}^{\infty} \frac{x^{3}}{\exp(x) - 1} dx = \frac{\pi^{4}}{15}$$

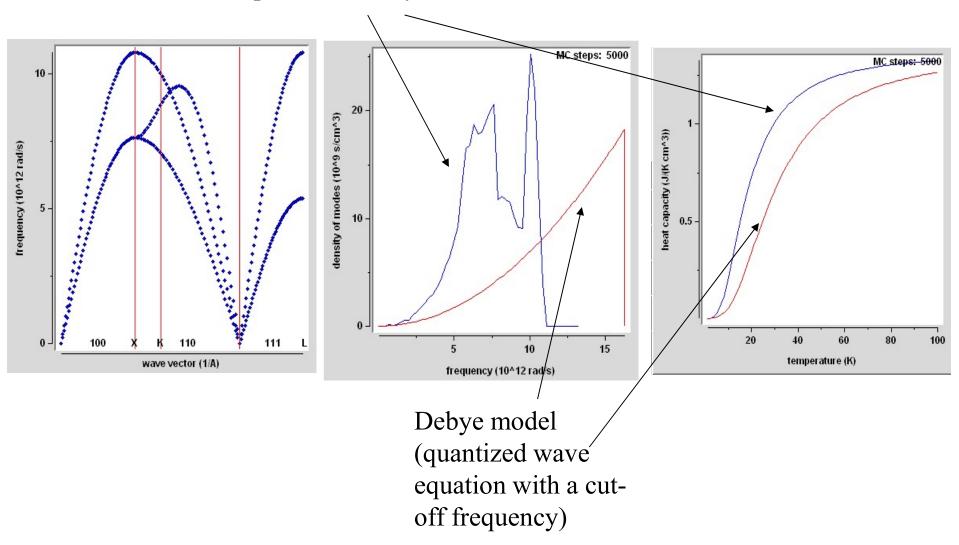
$$u \approx \frac{3\pi^4}{5} n k_B \frac{T^4}{\theta_D^3} \qquad c_v \approx \frac{12\pi^4}{5} n k_B \left(\frac{T}{\theta_D}\right)^3$$

Li	Be		Ta	ble 1	De	bye te	mpera	ture	and the	rmal	con	ducti	vity			В	С	I	V	0	F		Ne
344 0.85	1440 2.00															0.27	223	Marian Company					75
Na	Mg														ľ	AI	Si	F	)	s	С	1	Ar
158 1.41	400 1.56	1.2							it of θ, 300 K, i							428 2.37	645		; <b>,</b> ·				92
K	Ca	Sc	Ti	V	'	Cr	Mn	F	С	o	Ni		Cu	Zn		Ga	Ge	A	s	Se	В	r	Kr
91 1.02	230	360. 0.16	420 0.22		80 .31	630 0.94	410	47		45 .00	450 0.9		343 4.01	327 1.1		320 0.41	374 0.60		82 ).50	90 0.02			72
Rb	Sr	Υ	Zr	N	b	Мо	Тс	R	u R	h	Pd	1	Ag	Cd		In	Sn	" s	b	Те	Ti		Xe
56 0.58	147	280 0.17	291 0.23		75 .54	450 1.38	0.51	60		30 50	274 0.7		225 4.29	209	EDUCATION AND	108 0.82	200		11 .24	153 0.02			64
Cs	Ва	La β	Hf	Та	a	w	Re	0:	s Ir		Pt	1	۱u	Hg		TI	Pb	В	i	Ро	Ai	:	Rn
38 0.36	110	142 0.14	252 0.23		40 .58	400 1.74	430 0.48	50 0.		20 47	240 0.7		.65 3.17	71.		78.5 0.46	105 0.35		19 .08				
Fr	Ra	Ac	7		_																		
				Се	Pr	. N	d I	Pm	Sm	Eu		Gd	Tb		Dy	н	0	Er	Tm	, ,	Ύb	Lu	
			U	0.11	0.	12 0.	16		0.13			200 0.11	0.	11	210 0.1	1 0.	16	0.14	0.	CONTRACTOR DESIGNATION AND DESIGNATION OF THE PERSON NAMED IN CONTRACTOR O	120 0.35	210 0.1	ACCUPATION OF THE PARTY OF THE
				Th	Pa	U	ı	۷p	Pu	An	n	Cm	Bk		Cf	Es		Fm	Md	1	Vo	Lr	
				163 0.54		20 0.	)7 28 (	0.06	0.07			•											

Kittel

#### Phonon density of states

fcc phonon density of states



#### Thermal properties

internal energy density 
$$u = \int_{0}^{\infty} u(\omega)d\omega = \int_{0}^{\infty} \frac{\hbar\omega D(\omega)}{\exp\left(\frac{\hbar\omega}{k_{B}T}\right) - 1} d\omega$$
 [ J/m<sup>3</sup>]

specific heat 
$$c_{v} = \frac{du}{dT} = \int \left(\frac{\hbar\omega}{T}\right)^{2} \frac{D(\omega) \exp\left(\frac{\hbar\omega}{k_{B}T}\right)}{k_{B} \left(\exp\left(\frac{\hbar\omega}{k_{B}T}\right) - 1\right)^{2}} d\omega \quad [J \text{ K}^{-1} \text{ m}^{-3}]$$

entropy density 
$$s(T) = \int \frac{c_v}{T} dT = \frac{1}{T} \int_0^\infty \frac{\hbar \omega D(\omega)}{\exp\left(\frac{\hbar \omega}{k_B T}\right) - 1} d\omega \quad [J \text{ K}^{-1} \text{ m}^{-3}]$$

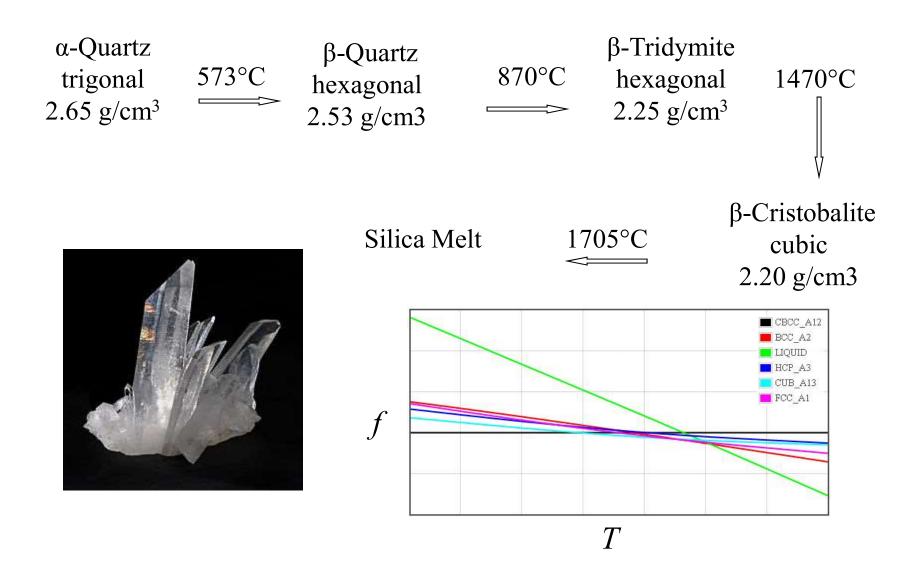
Helmholtz free energy density

$$f(T) = u - Ts = k_B T \int_0^\infty D(\omega) \ln \left( 1 - \exp\left(\frac{-\hbar \omega}{k_B T}\right) \right) d\omega \quad \left[ J/m^3 \right]$$

#### Phonons

	Linear Chain $m\frac{d^2u_s}{dt^2} = C(u_{s+1} - 2u_s + u_{s-1})$	Linear chain 2 masses $M_1 \frac{d^2 u_s}{dt^2} = C(v_{s-1} - 2u_s + v_s)$ $M_2 \frac{d^2 v_s}{dt^2} = C(u_s - 2v_s + u_{s+1})$	$\frac{d^2 u_{lmn}^x}{dt^2} = \frac{C}{\sqrt{3} \ m} [(u_{l+1m+1n+1}^x - u_{lmn}^x) + (u_{l-1m+1n+1}^x - u_{ln}^x) \\ + (u_{l+1m-1n+1}^x - u_{lmn}^x) + (u_{l+1m+1n-1}^x - u_{lmn}^x) + (u_{l-1m+1n+1}^x + u_{ln}^x) \\ + (u_{l+1m-1n-1}^x - u_{lmn}^x) + (u_{l+1m+1n+1}^x - u_{lmn}^y) - (u_{l-1m+1n+1}^y) \\ + (u_{l+1m-1n+1}^y - u_{lmn}^y) - (u_{l+1m+1n-1}^y - u_{lmn}^y) + (u_{l-1m+1n+1}^y) \\ + (u_{l+1m-1n-1}^y - u_{lmn}^y) + (u_{l+1m+1n+1}^y - u_{lmn}^y) - (u_{l-1m+1n+1}^y) \\ + (u_{l+1m-1n+1}^z - u_{lmn}^z) - (u_{l+1m+1n-1}^z - u_{lmn}^z) + (u_{l-1m+1n+1}^z - u_{lmn}^z) \\ - (u_{l+1m-1n-1}^z - u_{lmn}^z)]$ And similar expressions for the y and z of the part of the pa				
Eigenfunction solutions	$u_s = A_k e^{i(ksa-\alpha x)}$	$u_s = ue^{i(ksa-at)}$ $v_s = ve^{i(ksa-at)}$	$u_{lmn}^x = u_{\overrightarrow{k}}^x e^{i(l\overrightarrow{k} \cdot \overrightarrow{a_1} + m\overrightarrow{k} \cdot \overrightarrow{a_2} + n\overrightarrow{k} \cdot \overrightarrow{a_3})} = u_{\overrightarrow{k}}^x e^{i(\frac{(-l-1)^2}{2})}$ And similar expressions for the $y$ and $z$				
Dispersion relation	$\omega = \sqrt{\frac{4C}{m}} \left  \sin\left(\frac{ka}{2}\right) \right $ $\frac{a}{\sqrt{4C/m}}$ $\frac{a}{\sqrt{\frac{\pi}{a}}}$ $0$ $0$ $0$ $0$ $0$ $0$ $0$ $0$ $0$ $0$	$\omega^2 = C \left( \frac{1}{M_1} + \frac{1}{M_2} \right) \pm C \sqrt{\left( \frac{1}{M_1} + \frac{1}{M_2} \right)^2 - \frac{4 \sin^2 \left( \frac{ka}{2} \right)}{M_1 M_2}}$ $\left[ 2C \left( \frac{1}{M_1} + \frac{1}{M_2} \right) \right]^{1/2}$ $(2C/M_2)^{1/2}$ $(2C/M_1)^{1/2}$ $\frac{\pi}{a}$ Calculate $\omega(k)$	The dispersion $ \frac{4 - \cos(\frac{\alpha}{2}(k_x + k_y + k_z)) - \cos(\frac{\alpha}{2}(3k_x - k_y - k_z))}{-\cos(\frac{\alpha}{2}(-k_x + 3k_y - k_z)) - \cos(\frac{\alpha}{2}(-k_x + k_y + 3k_z)) - \frac{m\omega^2}{\sqrt{3}C}} + \cos(\frac{\alpha}{2}(-k_x + 3k_y - k_z)) + \cos(\frac{\alpha}{2}(3k_x - k_y - k_z)) + \cos(\frac{\alpha}{2}(k_x + k_y + k_z)) - \cos(\frac{\alpha}{2}(-k_x - k_y + 3k_z)) - \cos(\frac{\alpha}{2}(-k_x + 3k_y - k_z)) - \cos(\frac{\alpha}{2}(-k_x - k_y + 3k_z)) + \cos(\frac{\alpha}{2}(-k_x + k_y + k_z)) + \cos(\frac{\alpha}{2}(-k_x - k_y + 3k_z)) + \cos(\frac{\alpha}{2}(-k_x + k_y + k_z)) + \cos(\frac{\alpha}{2}(-k_x - k_y + 3k_z)) + \cos(\frac{\alpha}{2}(-k_x + k_y + k_z)) + \cos(\frac{\alpha}{2}(-k_x - k_y + 3k_z)) + \cos(\frac{\alpha}{2}(-k_x + k_y + k_z)) + \cos(\frac{\alpha}{2}(-k_x - k_y + 3k_z)) + \cos(\frac{\alpha}{2}(-k_x + k_y + k_z)) + \cos(\frac{\alpha}{2}(-k_x - k_y + 3k_z)) + \cos(\frac{\alpha}{2}(-k_x + k_y + k_z)) + \cos(\frac{\alpha}{2}(-k_x - k_y + 3k_z)) + $				
Density of states $D(k)$	$D(k) = \frac{1}{\pi}$	$D(k) = \frac{1}{\pi}$	$D(k) = \frac{3k^2}{2\pi^2}$				

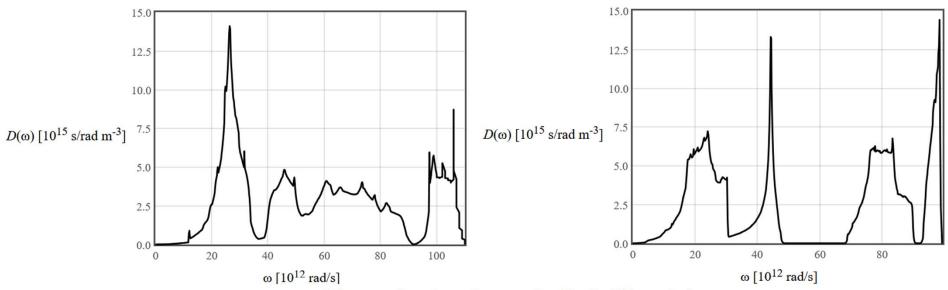
#### Quartz



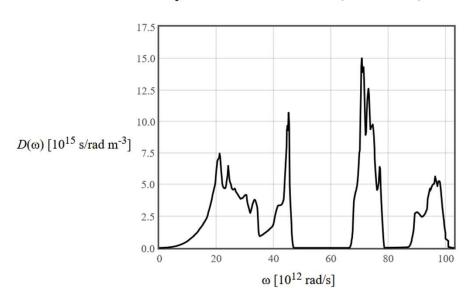
#### ZnO

#### Phonon density of states for ZnO (Rocksalt)

#### Phonon density of states for ZnO (Zincblende)



Phonon density of states for ZnO (Wurtzite)



#### Waves and particles

The eigen function solutions of the wave equation are plane waves. The scattering time is one over the rate for scattering from a given plane wave solution to any other.

Phonons are particles. The scattering time is the time before the phonons scatter and randomly change energy and momentum.

$$E = \hbar \omega$$
$$\vec{p} = \hbar \vec{k}$$

The average time between scattering events is  $\tau_{sc} = 1/\Gamma$ 

#### Phonon scattering

Scattering randomizes the momentum of the phonons.

$$H = H_{HO} + H_1$$

Transition rates determined by Fermi's golden rule

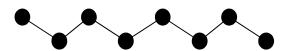
$$\Gamma_{i \to f} = \frac{2\pi}{\hbar} \left| \left\langle \psi_f \left| H_1 \middle| \psi_i \right\rangle \right|^2 \delta \left( E_f - E_i \right) \right|$$

Any process (3 phonon, 4 phonon, 5 phonon. ...) that conserves energy and momentum is allowed.

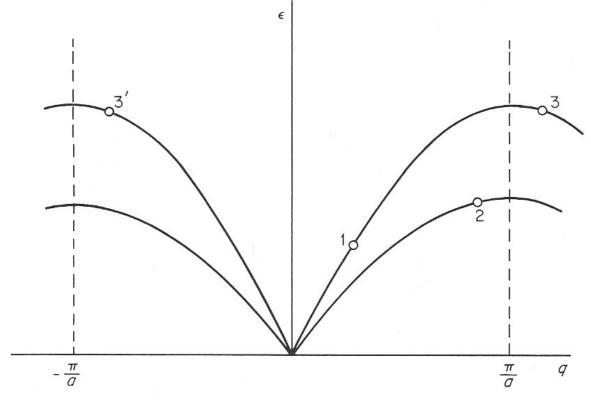
Results in attenuation of acoustic waves

#### **Umklapp Processes**

Three phonon scattering



$$\hbar \vec{k}_1 + \hbar \vec{k}_2 = \hbar \vec{k}_3 + \hbar \vec{G}$$

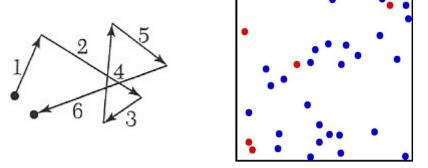


from: Hall, Solid State Physics

#### Heat transport (Kinetic theory)

Treat phonons as an ideal gas of particles that are confined to the volume of the solid.

Phonons move at the speed of sound. They scatter due to imperfections in the lattice and anharmonic terms in the Hamiltonian.



The average time between scattering events is  $\tau_{sc}$ 

The average distance traveled between scattering events is the mean free path:  $l = v\tau_{sc} \sim 10 \text{ nm}$ 

#### Diffusion equation/ heat equation

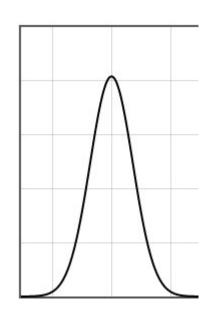
$$\frac{dn}{dt} = -D\nabla^2 n$$
Diffusion constant

Fick's law

$$\vec{j} = -D\nabla n$$

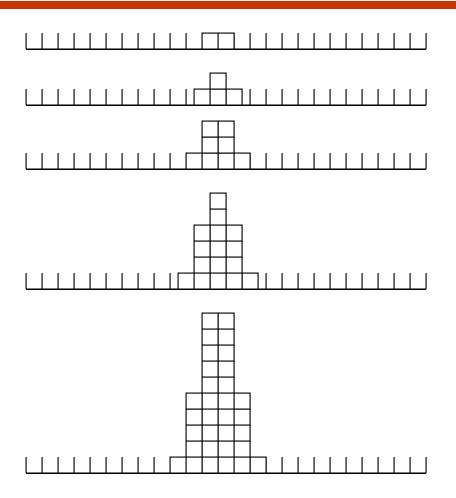
Continuity equation

$$\frac{dn}{dt} = \nabla \cdot \vec{j}$$

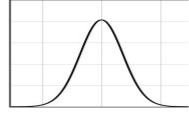


$$n = \frac{1}{\sqrt{4\pi Dt}} \exp\left(\frac{-r^2}{4Dt}\right)$$

#### Random walk



$$\frac{\Delta n_s}{\Delta t} = n_{s+1} - 2n_s + n_{s-1}$$



$$\exp(-x^2)$$

Central limit theorem: A function convolved with itself many times forms a Gaussian